Lectures on Brane World Scenarios

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Fine-tuning (naturalness) problem in the Standard Model

Standard model: gauge theory $SU(3)_c \times SU(2)_L \times U(1)_Y$ quarks & leptons & <u>Higgs</u>

Non-zero Higgs VEV determines the vacuum $SU(2)_L \times U(1)_Y \rightarrow U(1)_{em}$

W, Z bosons get masses of order 100 GeV

Serious problem: SM vacuum is not stable under quantum corrections

$$\Delta m_H^2 = H - - \Phi^{\dagger} \sim -\Lambda_{new}^2$$

$$\langle H \rangle \sim m_H^2 = (m_H^{tree})^2 + \Delta m_H^2$$

 $\sim (m_H^{tree})^2 - \Lambda_{new}^2$

If new physics scale >> electroweak scale

For example
$$\Lambda_{new} = M_{Planck} \sim 10^{19} \text{GeV}$$

 $\langle H \rangle = (m_H^{tree})^2 - \Lambda_{new}^2$
 $\rightarrow \mathcal{O}(M_W^2) = \mathcal{O}((10^{19} \text{GeV})^2) - \mathcal{O}((10^{19} \text{GeV})^2)$

Fine-tuning is needed!

In order to solve this problem

I) New Physics without quadratic divergence in quantum corrections



II) New Physics scale is close to the electorweak scale

$$\Lambda_{new} \sim \mathcal{O}(1 \text{TeV})$$

→ Brane World Scenario

Brane World Scenario

String Theory : classical solution (Polchinski '95) <u>D-brane</u>: D+1 dim. object embedded in higher dim. space time



Open string \rightarrow fermion, boson, gauge <u>on D-brane</u> Closed string \rightarrow graviton <u>in the bulk</u>

Phenomenological model



Beyond the standard model → Brane World Scenario 4+ dimensions New property "geometry"

 Typical Scenario:
 Large (flat) Extra Dimensions
 (Arkani-Hamed-Dimopoulos-Dvali, '98)

 Warped (small) Extra Dimensions (Randall-Sundrum, '99)

Large Extra Dimension Scenario

Solution to fine-tuning problem without SUSY

<u>Low scale gravity model in 4+</u> dimensions $\Lambda_{new} = M_{4+\delta} \sim \mathcal{O}(1 \text{TeV})$

$$S_{4+\delta} = \left(\frac{M_{4+\delta}}{M_{4+\delta}}\right)^{2+\delta} \int d^4x d^\delta y \sqrt{-g_{4+\delta}} R_{4+\delta}$$
$$= M_{Planck}^2 \int d^4x \sqrt{-g_4} R_4$$

 $M_{Planck}^2 = M_{4+\delta}^{2+\delta} V_{\delta}$ V : volume of extra-dim. $V_{\delta} = (2\pi R)^{\delta}$ (compactified on T^{δ})

For $M_{4+\delta} \sim 1 \text{TeV}$

$$\delta$$
 r 1 10^{13} cmexcluded2 10^{-1} mmallowed3 10^{-6} mmallowedHoyle et al., PRL 86 (2001) 1418

General picture of Brane World Scenario

Setup: 4+ dimensional model ← compact extra-dimensions some 3-branes





Toy model



Equation of motion

$$\partial_M \partial^M \phi = (\partial_\mu \partial^\mu - \sum_{i=1}^{\delta} \partial_{y_i}^2) \phi = 0$$

Plane wave: $\phi = \tilde{\phi}(x)e^{ik_1y_1}e^{ik_2y_2}...e^{ik_\delta y_\delta}$

Compact extra-dimensions on T^{δ}

→
$$\phi(x, y_1, y_2, ..., y_\delta) = \phi(x, y_1 + 2\pi R_1, ..., y_\delta + 2\pi R_\delta)$$

$$\begin{array}{l} \bullet \quad k_i = \frac{n_i}{R_i} \; ; \; n_i = 0, \pm 1, \pm 2, .. \\ \\ \partial_M \partial^M \phi \to \left(\partial_\mu \partial^\mu + \sum_{i=1}^{\delta} \left(\frac{n_i}{R_i} \right)^2 \right) \tilde{\phi}(x) = 0 \end{array}$$

Kaluza-Klein modes:

$$m_{n_1,n_2,\dots,n_{\delta}}^2 = \sum_{i=1}^{\delta} \left(\frac{n_i}{R_i}\right)^2$$

Example: 5(=4+1) dim. model (compact 5-th dim. on S^1)

$$\mathcal{L}_{brane} = \bar{\psi}(x)i\gamma^{\mu}\partial_{\mu}\psi(x) - g\bar{\psi}(x)\psi(x)\phi(x,y=0)$$
$$\mathcal{L}_{bulk} = \frac{1}{2}\left(\partial_{\mu}\phi\right)\left(\partial^{\mu}\phi\right) - \frac{1}{2}\left(\partial_{y}\phi\right)^{2}$$

KK mode expansion:

$$\phi(x,y) = \phi_{odd}(x,y) + \phi_{even}(x,y)$$

$$\phi_{odd}(x,y) = \sum_{n=1}^{\infty} \phi_{odd}^{(n)}(x) \sin\left(\frac{n}{R}y\right)$$

$$\phi_{even}(x,y) = \phi_{even}^{(0)}(x) + \sum_{n=1}^{\infty} \phi_{even}^{(n)}(x) \cos\left(\frac{n}{R}y\right)$$

<u>4 dimensional description</u>

$$\mathcal{L} = \int_{0}^{2\pi R} dy \mathcal{L}_{bulk}$$

= $\frac{1}{2} (\pi R) \sum_{n=1} \left[(\partial_{\mu} \phi(x)_{odd}^{(n)})^2 - \left(\frac{n}{R}\right)^2 (\phi(x)_{odd}^{(n)})^2 \right]$
+ $\frac{1}{2} (2\pi R) (\partial_{\mu} \phi(x)_{even}^{(0)})^2$
+ $\frac{1}{2} (\pi R) \sum_{n=1} \left[(\partial_{\mu} \phi(x)_{even}^{(n)})^2 - \left(\frac{n}{R}\right)^2 (\phi(x)_{even}^{(n)})^2 \right]$

$$\mathcal{L}_{brane} = \bar{\psi} i \gamma^{\mu} \partial_{\mu} \psi - g \bar{\psi} \psi \left(\phi_{even}^{(0)} + \sum_{n=1} \phi_{even}^{(n)} \right)$$

After rescaling

$$\mathcal{L} = \frac{1}{2} \sum_{n=1}^{\infty} \left[(\partial_{\mu} \phi(x)_{odd}^{(n)})^2 - \left(\frac{n}{R}\right)^2 (\phi(x)_{odd}^{(n)})^2 \right] \\ + \frac{1}{2} (\partial_{\mu} \phi(x)_{even}^{(0)})^2 \\ + \frac{1}{2} \sum_{n=1}^{\infty} \left[(\partial_{\mu} \phi(x)_{even}^{(n)})^2 - \left(\frac{n}{R}\right)^2 (\phi(x)_{even}^{(n)})^2 \right]$$

$$\mathcal{L}_{brane} = \bar{\psi} i \gamma^{\mu} \partial_{\mu} \psi$$
$$- \bar{\psi} \psi \left(\frac{g}{\sqrt{2\pi R}} \phi_{even}^{(0)} + \sum_{n=1}^{\infty} \frac{g}{\sqrt{\pi R}} \phi_{even}^{(n)} \right)$$

Characteristic features: KK mode mass n/R universal coupling

Large Extra Dimension Scenario

Phenomenology

 $\cdot \mathbf{M}_{4+} \sim \mathcal{O}(1 \mathrm{TeV})$

'many graviton Kaluza-Klein modes

$$\begin{aligned} G_{\mu\nu}(x^{\mu}, y^{1}, y^{2}, ..., y^{n}) &= \sum_{n} g_{\mu\nu}^{(\vec{n})}(x^{\mu}) \ \chi^{(\vec{n})}(\vec{y}) \\ \chi^{(n)} \propto e^{i\frac{\vec{n}\cdot\vec{y}}{r}} \ , \ \left(m_{KK}^{(\vec{n})}\right)^{2} &= \frac{|\vec{n}|^{2}}{r^{2}} \quad \text{If 6 dim.} \Rightarrow \quad \frac{1}{r} \sim 10^{-4} \text{eV} \end{aligned}$$

Phenomenology of Extra-dimension Scenario

= Phenomenology of graviton Kaluza-Klein modes

Construction of effective action in 4D

<u>1.4D Lagrangian</u>

4+ dim. Graviton → 4D graviton in 4D KK gravitons KK gravi-scalars KK gravi-vectors

2. Feynman rules

<u>4D reduction of 4+</u> dim. Einstein's Eqs.</u>

$$\Box_{4+\delta}G_{MN} = -\frac{T_{MN}}{(M_{4+\delta})^{2+\delta}}$$

$$G_{MN} = \begin{bmatrix} G_{\mu\nu}^{(n)} & V_{\mu j}^{(n)} \\ V_{\mu i}^{(n)} & H^{(n)}\delta_{ij} + S_{ij}^{(n)} \end{bmatrix}$$
KK graviton
$$(\Box + \hat{n})G_{\mu\nu}^{(n)} = \frac{1}{\bar{M}_P} \left[-T_{\mu\nu} + \frac{1}{3} \left(\frac{\partial_{\mu}\partial_{\nu}}{\hat{n}^2} \right) T_{\lambda}^{\lambda} \right]$$
KK gravi-scalar
$$(\Box + \hat{n})H^{(n)} = \frac{\sqrt{\frac{3(\delta-1)}{\delta+2}}}{\bar{M}_P} T_{\mu}^{\mu}$$
Vector
$$(\Box + \hat{n})V_{\mu j}^{(n)} = 0$$
Scalar
$$(\Box + \hat{n})S_{ij}^{(n)} = 0$$

Feynman rules

$$T^{\mu\nu} = 2 \frac{\partial \mathcal{L}}{\partial g_{\mu\nu}} - g^{\mu\nu} \mathcal{L}$$

ex) QED + KK graviton

$$\frac{f(k_1)}{f(k_2)} \longrightarrow G_{\mu\nu}^{(n)} - \frac{i}{4\bar{M}_P} [(k_1 + k_2)_{\mu}\gamma_{\nu} + (\nu \leftrightarrow \mu)]$$

$$\begin{array}{c} A_{\alpha}(k_{1}) \\ A_{\beta}(k_{2}) \end{array} \qquad G_{\mu\nu}^{(n)} \qquad -\frac{1}{M_{P}} \quad \left[\begin{array}{c} \frac{1}{2} \eta_{\mu\nu}(k_{1\beta}k_{2\alpha} - (k_{1}k_{2})\eta_{\alpha\beta}) + \eta_{\alpha\beta}k_{1\mu}k_{2\nu} \\ + \eta_{\mu\alpha}(\eta_{\nu\beta}(k_{1}k_{2}) - k_{1\beta}k_{2\nu}) - \eta_{\mu\beta}k_{1\nu}k_{2\alpha} \\ + (\mu \leftrightarrow \nu) \right] \end{array}$$



$$-\frac{ieQ}{2\bar{M}_P}[\gamma_\mu\eta_{\nu\alpha}+(\mu\leftrightarrow\nu)]$$

Detection of Extra-dimension @ <u>future colliders</u>

→ detection of KK graviton

direct → KK graviton emission indirect → KK graviton mediated process

First detection of <u>spin 2 particle</u> !

KK graviton emission process

Emitted KK graviton → non-interacting & stable particle → missing energy event

Example: $e^+e^- \rightarrow \gamma + \text{nothing}$

SM background:
$$e^+e^- \rightarrow \gamma \bar{\nu} \nu$$



Each process is suppressed by $1/\bar{M_P}^2$ <u>But</u> × of KK modes $\Longrightarrow \frac{1}{M_{4+\delta}^2}$

<u>**Result</u>** $e^+e^- \rightarrow \gamma + E_{\text{missing}}$ </u>



$$\sigma \propto \left(\frac{\sqrt{s}}{M_{4+\delta}}\right)^{\delta+2}$$

Virtual KK graviton mediated process



Need regularization

Naïve: Cut Off by $m_{KK}^{MAX} \sim M_{4+\delta}$

$$\frac{4\pi\lambda}{M_S^4} = -\frac{8\pi}{M_P^2} \sum_{\vec{n}} \frac{1}{s - \left(m_{KK}^{(\vec{n})}\right)^2}; \qquad \lambda = \pm 1$$

$$\mathcal{M} = \frac{4\pi\lambda}{M_S^4} T_{\mu\nu}(p_1, p_2) T^{\mu\nu}(p_3, p_4)$$

Calculations of some processes

$$e^+e^- \rightarrow \begin{array}{c} f\overline{f} \\ \gamma\gamma, W^+W^-, ZZ \\ hh \end{array}$$

$$\sum_{spin} |\mathcal{M}|^2 = \sum |\mathcal{M}_{SM} + \mathcal{M}_G|^2 = \sum \left(|\mathcal{M}_{SM}|^2 + \mathcal{M}_{IF} + |\mathcal{M}_G|^2 \right)$$

Interference:
$$\sum \mathcal{M}_{IF} = \sum \mathcal{M}_{SM}^{\dagger} \mathcal{M}_G + h.c. \propto \lambda$$

Only KK graviton: $\sum |\mathcal{M}_G|^2 \propto \lambda^2$

$$\mathcal{M} = \frac{4\pi\lambda}{M_S^4} T_{\mu\nu}(p_1, p_2) T^{\mu\nu}(p_3, p_4)$$

Results I: total cross section

 \leftarrow new physics evidence

II: angular dependence of cross section

← effects due to spin 2 particle exchange

$$e^+e^- \to f\bar{f}$$
 processes



$$e^+e^- \rightarrow \tau \bar{\tau}$$



Angular dependence of cross section

$$\frac{d\sigma(e^+e^- \to \tau\bar{\tau})}{d\cos\theta} \text{[fb]} \quad \text{at} \quad \sqrt{s} = 1 \text{TeV}$$



 $M_S = \infty$

$$e^+e^- \rightarrow t\bar{t}$$

$$\lambda = +1$$



 \sqrt{s} [GeV]

Angular dependence of cross section

$$\frac{d\sigma(e^+e^- \to tt)}{d\cos\theta}$$
[fb] at $\sqrt{s} = 1 \text{TeV}$







$$M_S = 2 \text{TeV}$$

 $M_S = 3 \text{TeV}$
 $M_S = 4 \text{TeV}$

$$\lambda = -1$$



$$m_t = 175 \text{GeV}$$

 $M_S = \infty$

 $e^+e^- \to VV$ processes



Angular distribution

$$\frac{d\sigma}{d\cos\theta}/\frac{d\sigma}{d\cos\theta}|_{SM}$$

$$\sqrt{s} = 1 {\rm TeV}$$

 $\lambda = +1$







 $M_S = 2 \text{TeV}$ $M_S = 3 \text{TeV}$ $M_S = 4 \text{TeV}$

$$e^+e^- \rightarrow h\bar{h}$$
 process

KK graviton exchange is dominant

SM background free \rightarrow very interesting

if this cross section is large enough

$$\sigma(e^+e^- \to hh) = \frac{\pi\lambda^2}{480M_S^8} \sqrt{1 - 4\frac{m_h^2}{s}} \left(s^3 - 8m_h^2s^2 + 16m_h^4s\right)$$



Angular dependence of cross section



$$\sum |\mathcal{M}|^2 = \frac{1}{2} \left(\frac{4\pi\lambda}{M_S^4}\right)^2 (t-u)^2 (tu-m_h^4)$$

Results of simulation studies (VERY Preliminary!)

 $e^+e^- \rightarrow hh$ process → $hh \rightarrow b\bar{b}b\bar{b}$ 4 jet events

$$\sqrt{s} = 1 \text{TeV}$$

 $M_S = 2 \text{TeV}$
 $m_h = 120 \text{GeV}$

Not including SM background



